

The CHY Formalism for Massless Scattering

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Overview

- 1 Scattering Amplitudes
- 2 scattering Equations
- 3 CHY formalism

1 Scattering Amplitudes

2 scattering Equations

3 CHY formalism

Feynman Diagrams

- For **QED** process, Feynman diagram is an efficient tool to calculate scattering amplitudes.

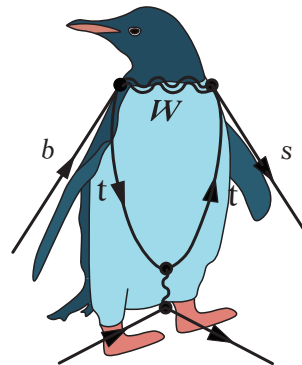


Figure: Kawaii feynman diagram

Feynman Diagrams

- For **QED** process, Feynman diagram is a efficient tool to calculate scattering amplitudes.
- Feynman's rule is derived from the Lagrangian, there are many terms in the Lagrangian that are blame to **gauge redundancy**, which inevitably leads to very complex expressions for individual Feynman diagrams, but the superposition of all Feynman diagrams is simple.

$$S_{\text{EH}} = \int d^D x \left[h \partial^2 h + \kappa h^2 \partial^2 h + \kappa^2 h^3 \partial^2 h + \kappa^3 h^4 \partial^2 h + \dots \right]$$

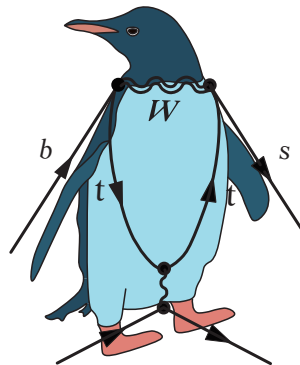


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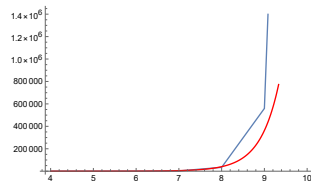


Figure: Too many diagrams

$$S_{\text{EH}} = \int d^D x \left[h \partial^2 h + \kappa h^2 \partial^2 h + \kappa^2 h^3 \partial^2 h + \kappa^3 h^4 \partial^2 h + \dots \right] \text{to sum}$$

- We have too many Feynman diagrams to sum. The number of diagrams is growing much **faster** than $n!$.

Chinese Magic

Spinor-Helicity Formalism

- Using On-shell condition $p^2 = 0$:

$$p_{a\dot{a}} = \sigma_{a\dot{a}}^\mu p_\mu = \lambda_a \tilde{\lambda}_{\dot{a}} \equiv |p] \langle p|$$

we can do the same thing for $\bar{\sigma}_{\dot{a}a}^\mu p_\mu$ to define $|p\rangle$ and $[p|$.

- It is easy to see that $|\cdot\rangle$ and $|\cdot]$ automatically satisfy the gauge invariant condition:

$$p^{\dot{a}b}|p]_b = 0, \quad p_{a\dot{b}}|p\rangle^{\dot{b}} = 0, \quad [p|^b p_{b\dot{a}} = 0, \quad \langle p|_{\dot{b}} p^{\dot{b}a} = 0.$$

- In fact we can use gauge degree of freedom to simplify our expressions:

$$\epsilon_\mu^+(k; q) = \frac{\langle q^- | \gamma_\mu | k^- \rangle}{\sqrt{2} \langle qk \rangle}$$

The reference momenta q can be chosen freely.

Tree Amplitudes of YM

- MHV Amplitudes (Park-Taylor):

$$A_n [1^+ \dots i^- \dots j^- \dots n^+] = \frac{\langle ij \rangle^4}{\langle 12 \rangle \langle 23 \rangle \dots \langle n1 \rangle}$$

- N^k MHV Amplitudes:

$$\begin{aligned} A_n^{\text{NPMHV}}(c_0, c_1, \dots, c_p, n) &= \frac{\delta^{(4)}(p)}{\langle 12 \rangle \langle 23 \rangle \dots \langle n1 \rangle} \times \\ &\times \sum_{\text{all paths of length } p} 1 \cdot \tilde{R}_{n; a_1 b_1} \cdot \tilde{R}_{n; \{l_2\}; \{a_2 b_2\}}^{\{L_2\}; \{U_2\}} \cdot \dots \cdot \tilde{R}_{n; \{l_p\}; \{a_p b_p\}}^{\{L_p\}; \{U_p\}} \\ &\times \left(\det \Xi_n^{\text{path}}(c_0, \dots, c_p) \right)^4 \end{aligned}$$

Complex, but fully solvable by computers

1 Scattering Amplitudes

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Riemann Sphere

Momentum Space

$$\mathfrak{K}_{D,n} := \{(k_1^\mu, k_2^\mu, \dots, k_n^\mu) \mid \sum_{a=1}^n k_a^\mu = 0, k_1^2 = k_2^2 = \dots = k_n^2 = 0\} / SO(1, D-1) \quad (1)$$

If there is no codimensional singularity

$$s_{a_1, a_2, \dots, a_r} := (k_{a_1} + k_{a_2} + \dots + k_{a_r})^2 \neq 0, \quad \forall r = 1, \dots, n \quad (2)$$

We can consider the moduli space of Riemann spheres \mathbb{CP}^1 with n distinct punctures on it to carve $\mathfrak{K}_{D,n}$ equivalently.

$$\mathfrak{M}_{0,n} \equiv \{\sigma_1, \sigma_2, \dots, \sigma_n\} / SL(2, \mathbb{C})$$

$$\mathfrak{K}_{D,n} \iff \mathfrak{K}_{D,n} \text{ by } k_a^\mu = \frac{1}{2\pi i} \oint_{|z-\sigma_a|=\epsilon} dz \frac{p^\mu(z)}{\prod_{b=1}^n (z-\sigma_b)} \quad (3)$$

Scattering Equations

Using eq.3, we can derive the scattering equation:

$$\sum_{b \neq a} \frac{s_{ab}}{\sigma_a - \sigma_b}, \quad a = 1, 2, \dots, n, \quad s_{ab} = 2k_a \cdot k_b \quad (4)$$

- n equations but only $n - 3$ of them are independent.

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- KLT orthogonality of solutions to scattering equations.

$$\frac{(i, j)}{(i, i)^{\frac{1}{2}}(j, j)^{\frac{1}{2}}} = \delta_{ij} \quad (5)$$

Where,

$$(i, j) := \sum_{\alpha, \beta \in \mathcal{S}_{n-3}} V^{(i)}(\alpha) S[\alpha|\beta] U^{(j)}(\beta)$$

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Where,

$$V(\omega) = \frac{1}{(\sigma_1 - \sigma_{\omega(2)})(\sigma_{\omega(2)} - \sigma_{\omega(3)}) \cdots (\sigma_{\omega(n-2)} - \sigma_{n-1})(\sigma_{n-1} - \sigma_n)(\sigma_n - \sigma_1)},$$

$$U(\omega) = \frac{1}{(\sigma_1 - \sigma_{\omega(2)})(\sigma_{\omega(2)} - \sigma_{\omega(3)}) \cdots (\sigma_{\omega(n-2)} - \sigma_n)(\sigma_n - \sigma_{n-1})(\sigma_{n-1} - \sigma_1)}.$$

1 Scattering Amplitudes

2 scattering Equations

3 CHY formalism

Bi-adjoint Scalar, YM and Einstein Gravity

The amplitudes of many QFTs (at the tree level) can be expressed by a unified formula (Cachazo, He, Yuan, 2013).

$$\mathcal{A}_n = \int d\mu_n \mathcal{I}_n, \quad d\mu_n = \frac{d^n \sigma}{\text{vol SL}(2, \mathbb{C})} \prod'_a \delta \left(\sum_{b \neq a} \frac{s_{ab}}{\sigma_{ab}} \right) \quad (6)$$

For different theories, the integral measure is the same, differing only in the CHY integrands \mathcal{I}_n

$$\mathcal{M}_n^{(s)} = \int \frac{d^n \sigma}{\text{vol SL}(2, \mathbb{C})} \prod'_a \delta \left(\sum_{b \neq a} \frac{s_{ab}}{\sigma_a - \sigma_b} \right) \left(\frac{\text{Tr}(T^{a_1} T^{a_2} \dots T^{a_n})}{(\sigma_1 - \sigma_2) \dots (\sigma_n - \sigma_1)} + \dots \right)^{2-s} (\text{Pf}' \Psi)^s \quad (7)$$

For Scalar, $s = 0$,

$$\mathcal{L}^{\Phi^3} := -\frac{1}{2} \partial_\mu \Phi_{I, \tilde{I}} \partial^\mu \Phi^{I, \tilde{I}} - \frac{\lambda}{3!} f_{I, J, K} \tilde{f}_{\tilde{I}, \tilde{J}, \tilde{K}} \Phi^{I, \tilde{I}} \Phi^{J, \tilde{J}} \Phi^{K, \tilde{K}} \quad (8)$$

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For Scalar, $s = 0$,

$$\mathcal{I}_{U(N) \times U(\tilde{N})}^{\Phi^3} := \mathcal{C}_{U(N)} \mathcal{C}_{U(\tilde{N})} \quad (8)$$

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Where

$$\mathcal{C}_{U(N)} := \sum_{\alpha \in S_n / \mathbb{Z}_n} \text{tr}(T^{\alpha(1)} T^{\alpha(2)} \dots T^{\alpha(n)}) \text{PT}_n(\alpha)$$

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Where,

$$\text{PT}_n[\alpha] := \frac{1}{\sigma_{\alpha(1)\alpha(2)} \sigma_{\alpha(2),\alpha(3)} \dots \sigma_{\alpha(n),\alpha(1)}}$$

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Delta functions in eq.6 totally local the integral, so we don't need to calculate annoying integral, we just need to solve the scattering equations.

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i.e.

$$\sum_{\{\sigma\} \in \text{solutions}} \frac{(\sigma_{pq} \sigma_{qr} \sigma_{rp})(\sigma_{ij} \sigma_{jk} \sigma_{ki})}{|\Phi|_{pqr}^{ijk}} \mathcal{I} \quad (8)$$

Bi-adjoint Scalar, YM and Einstein Gravity

For YM, $s = 1$, (*Gervais–Neveu gauge, ghost free*)

$$\mathcal{L} = \text{Tr} \left(-\frac{1}{2} \partial_\mu A_\nu \partial^\mu A^\nu - i\sqrt{2}g \partial^\mu A^\nu A_\nu A_\mu + \frac{g^2}{4} A^\mu A^\nu A_\nu A_\mu \right)$$

$$\mathcal{I}_n^{\text{YM}} = \mathcal{C}_n \text{Pf}' \Psi(\{k, \epsilon, \sigma\}) \quad (9)$$

For Gravitons, $s = 2$, (*de-Donder gauge, ghost free*)

$$\mathcal{L}_{\text{EH}} = \partial_\alpha h \partial_\beta h^{\alpha\beta} - \partial_\alpha h_{\beta\gamma} \partial^\beta h^{\alpha\gamma} - \frac{1}{2} (\partial_\alpha h)^2 + \frac{1}{2} (\partial_\gamma h_{\alpha\beta})^2 + \mathcal{O}(\kappa, h^3)$$

$$+ \partial^\nu h_{\mu\nu} \partial^\rho h_\rho^\mu + \frac{1}{4} (\partial_\mu h)^2 - \partial^\nu h_{\mu\nu} \partial^\mu h$$

$$\mathcal{I}_n^{\text{GR}} = \text{Pf}' \Psi_n \text{Pf}' \tilde{\Psi}_n \quad (10)$$

We introduce $\tilde{\Psi}$, because generally, we can contain dilatons and B-fields in GR. For pure graviton scattering, $\Psi = \tilde{\Psi}$.

Bi-adjoint Scalar, YM and Einstein Gravity

For YM, $s = 1$,

$$\mathcal{I}_n^{\text{YM}} = \mathcal{C}_n \text{Pf}' \Psi(\{k, \epsilon, \sigma\}) \quad (9)$$

For Gravitons, $s = 2$,

$$\mathcal{I}_n^{\text{GR}} = \text{Pf}' \Psi_n \text{Pf}' \tilde{\Psi}_n \quad (10)$$

Where,

$$\Psi = \begin{pmatrix} A & -C^T \\ C & B \end{pmatrix}, \quad \text{Pf}' \Psi := \frac{(-1)^{i+j}}{(\sigma_i - \sigma_j)} \text{Pf}(\Psi_{ij}^{ij}) \quad (11)$$

and,

$$A_{ab} = \begin{cases} \frac{s_{ab}}{\sigma_a - \sigma_b}, & a \neq b \\ 0, & a = b, \end{cases} \quad B_{ab} = \begin{cases} \frac{2\epsilon_a \cdot \epsilon_b}{\sigma_a - \sigma_b}, & a \neq b \\ 0, & a = b, \end{cases} \quad C_{ab} = \begin{cases} \frac{2\epsilon_a \cdot k_b}{\sigma_a - \sigma_b}, & a \neq b \\ -\sum_{c \neq a} \frac{2\epsilon_a \cdot k_c}{\sigma_a - \sigma_c}, & a = b \end{cases} \quad (12)$$

More theories and their connections

- The greatest advance in scattering amplitudes in the last two decades has been the formulation of the BCFW recursion relations, which has greatly simplified the calculations.

$$A_n = \sum_{\text{diagrams } I} \hat{A}_L(z_I) \frac{1}{P_I^2} \hat{A}_R(z_I)$$
$$= \sum_{\text{diagrams } I} \hat{i} \text{ --- } \text{L} \text{ --- } \hat{P}_I \text{ --- } \text{R} \text{ --- } \hat{j}$$

Figure: BCFW Recursion Relations.

More theories and their connections

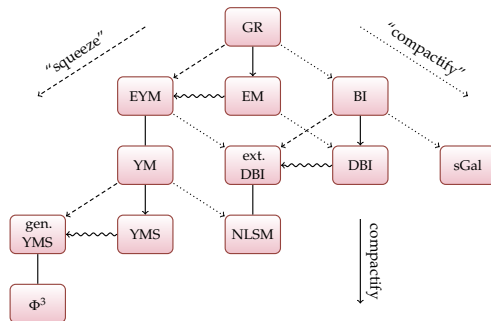
- Scattering equations are very difficult to find all $(n-3)!$ solutions. In this sense, the CHY formula doesn't bring us a new efficient tool for calculating amplitudes. But it gives us a unified framework to consider connections between different theories.

Theory	Integrand	Section
Einstein gravity	$\text{Pf}'\Psi_n \text{Pf}'\Psi_n$	4.5
Yang–Mills	$C_n \text{Pf}'\Psi_n$	4.4.1
Φ^3 flavored in $U(N) \times U(\tilde{N})$	$C_n C_n$	4.2.1
Einstein–Maxwell	$\text{Pf}[\mathcal{X}_n]_\gamma \text{Pf}'[\Psi_n]_{\tilde{\gamma}} \text{Pf}'\Psi_n$	5.1.3
Einstein–Yang–Mills	$C_{\text{tr}_1} \cdots C_{\text{tr}_t} \text{Pf}'\Pi(h; \text{tr}_1 \dots, \text{tr}_t) \text{Pf}'\Psi_n$	5.2
Yang–Mills–Scalar	$C_n \text{Pf}[\mathcal{X}_n]_s \text{Pf}'[\Psi_n]_{\tilde{s}} \text{Pf}'\Psi_n$	5.1.1
generalized Yang–Mills–Scalar	$C_n C_{\text{tr}_1} \cdots C_{\text{tr}_t} \text{Pf}'\Pi(g; \text{tr}_1 \dots, \text{tr}_t)$	5.2.4
Born–Infeld	$\text{Pf}'\Psi_n (\text{Pf}'A_n)^2$	4.4.3
Dirac–Born–Infeld	$\text{Pf}[\mathcal{X}_n]_s \text{Pf}'[\Psi_n]_{\tilde{s}} (\text{Pf}'A_n)^2$	5.1.2
extended Dirac–Born–Infeld	$C_{\text{tr}_1} \cdots C_{\text{tr}_t} \text{Pf}'\Pi(\gamma; \text{tr}_1 \dots, \text{tr}_t) (\text{Pf}'A_n)^2$	5.2.5
$U(N)$ non-linear sigma model	$C_n (\text{Pf}'A_n)^2$	4.2.3
special Galileon	$(\text{Pf}'A_n)^4$	4.2.6

Figure: CHY integrands of different QFTs

More theories and their connections

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Connections among integrands. Compactify: →. Squeeze: --->. "Compactify":>. Non-Abelian: ~~~~. Restrict to single trace: —.

Figure: Theory Web.

KLT double copy

$$\text{Gravity} = \text{YM}^2 / \phi^3,$$

$$M_n = (-1)^n \sum_{\beta, \gamma} \frac{A_n(1, \beta_{2,n-1}, n) \tilde{\mathcal{S}}[\beta_{2,n-1} | \gamma_{2,n-1}]_{p_n} \tilde{A}_n(n, \gamma_{2,n-1}, 1)}{s_{23\dots n}} \quad (13)$$

Where, KLT kernel $\mathcal{S}[\alpha|\beta]$ can be constructed by double partial amplitudes of ϕ^3 theory:

$$S[\alpha|\beta] = \prod_{i=2}^{n-2} \left(s_{1,\alpha(i)} + \sum_{j=2}^{i-1} \theta(\alpha(j), \alpha(i))_{\beta} s_{\alpha(j), \alpha(i)} \right) = m(\alpha|\beta)^{-1} \quad (14)$$

So, we divide YM^2 by ϕ^3 .

KLT relations

KLT orthogonality eq.5 can be rewritten as,

$$\delta_{\alpha,\gamma} = \sum_{\beta \in S_{n-3}} \int d\mu_n \mathbf{PT}(\alpha) \mathbf{PT}(\beta) S[\beta|\gamma] \quad (15)$$

So if we have a theory \mathcal{M}_n whose CHY integrand is $\mathcal{I} = \mathcal{I}^L \mathcal{I}^R$. Then we can define two partial amplitudes \mathcal{M}_n^L and \mathcal{M}_n^R , whose CHY integrands are $\mathcal{I}^L \cdot \mathbf{PT}$ and $\mathcal{I}^R \cdot \mathbf{PT}$, respectively.

Which gives a general KLT relation,¹

$$\mathcal{M}_n = \mathcal{M}_n^L \otimes_{\text{KLT}} \mathcal{M}_n^R \quad (16)$$

The CHY formalism enables us to gain insights that would otherwise be difficult to discern from the Lagrangian.

¹"KLT" suffix means we need a KLT kernel.

